Ray optics of generalised lenses (glenses)

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We study the ray optics of glenses, ideal thin lenses generalised to have different object- and image-sided focal lengths, and the most general light-ray-direction-changing surfaces that stigmatically image any point in object space to a corresponding point in image space. Gabor superlenses (Gabor, UK Patent 541,753 (1940); Hembd-Sölner et al., J. Opt. A: Pure Appl. Opt. 1, 94 (1999)) can be seen as pixellated realisations of glenses. Our analysis is centred on the nodal point. Whereas the nodal point of a thin lens always lies in the lens plane, that of a glens can lie anywhere on the optical axis. Utilising the nodal point, we derive simple equations that describe the mapping between object and image space and the light-ray-direction change. We demonstrate our findings with the help of ray-tracing simulations. Glenses allow novel optical instruments to be realised, at least theoretically, and our results facilitate the construction and analysis of such devices. © 2016 Optical Society of America

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1. INTRODUCTION

The thin-lens approximation is an idealised description of ray-optical imaging with lenses. The ideal thin lens is considered planar and – as the name suggests – of zero thickness, and any two or more light rays that intersect at a point (the object) before transmission through the lens (“in object space”) intersect again at a point (the image) after transmission through the lens (“in image space”). In each case, the definition of light ray includes not only the straight-line segment light travels before intersecting the lens, but also the straight-line continuation on either side.

Previously, a generalisation of ideal thin lenses was considered: the most general planar surface that redirects light rays without offsetting them — such that any point in object space is stigmatically imaged into a corresponding point in image space (and vice versa) is an ideal thin lens with different object- and image-sided focal lengths [1] (Fig. 1(a)). Note that there is no non-planar surface that can do this [2], so an ideal thin lens with two focal lengths is the most general surface of any shape that images in this way.

Thin lenses with different object- and image-sided focal lengths can be realised experimentally, with certain shortcomings, in the form of Gabor superlenses [3] (section 2), which comprise two confocal microlens arrays with different pitch. Gabor superlenses do not perform stigmatic imaging, in which the individual rays pass through the object and image position, but integral imaging, in which the axes of bundles of light rays intersect. Gabor superlenses have been demonstrated experimentally, and their imaging properties investigated, in Ref. [4].

Here we define an ideal thin lens with two focal lengths as a glens, which, according to taste, stands for generalised lens or Gabor superlens. The singular term glens should not be confused with the plural of “glen”, meaning a glacial valley in Scotland (Fig. 1(b)). Unlike earlier work, we study the nodal points of a glens, two of the cardinal points that describe the characteristics of an imaging system, and show that their positions coincide in a single nodal point (section 3). Utilising this nodal point, we derive vector expressions for the mapping between object and image space (section 4) and for the light-ray-direction change on transmission (section 5). We visualise the view through a few glenses using raytracing simulations (section 6).

2. PIXELLATED REALISATIONS OF GLENSES

For wave-optical reasons, a glens (or indeed an idealised thin lens) operating in air cannot be realised such that it correctly transforms all incident light-ray fields. This follows from considerations of optical path length [5], or by showing that the required light-ray-direction change can result in light-ray fields without wave-optical analog [6]. The latter argument relies on assuming that a corresponding phase front exists, and then showing that this cannot be the case as it would have to be discontinuous at every point.
The term integral image has its origin in integral photography, a technique invented more than a century ago [9] in which an image of a scene is taken not through a single lens, like in a normal camera, but through a planar lenslet (or microlens) array. In the simplest case the film (or image detector) is positioned in the focal plane of the lenslet array. A light-field camera (or plenoptic camera) [10, 11] is a modern version of just such a device, and algorithms have been invented that allow the captured light fields to be viewed in novel ways [12]. As each lenslet is in a different position, the resulting integral photograph represents the scene from a range of perspectives, and thus contains depth information about the scene. If the captured image is placed in the focal plane of an identical lenslet array and viewed through this lenslet array, a three-dimensional integral image of the scene appears.

The same integral image can be obtained without recording and developing the integral photograph, but by directly viewing the light distribution created by the first lenslet array in its focal plane with a second, identical, lenslet array that shares this focal plane, that is by simply viewing the scene through two identical lenslet arrays, separated by the sum of their focal lengths [13]. Such a pair of confocal lenslet arrays can thus be seen as an integral camera (the first lenslet array) and its viewer (the second lenslet array), with the integral photograph being formed in the common focal plane, but never recorded and developed, and instead viewed directly.

The individual telescopes can be generalised in a number of ways, for example by making the focal lengths of the two lenslet arrays different [14] or by shifting displacing the lenslet arrays relative to each other such that they remain confocal but such that the optical axes of the lenslets in one array no longer align with those of the lenslets in the other [15], resulting in generalised confocal lenslet arrays (GCLAs). The mapping between object space and (integral) image space can be so general that homogeneous GCLAs (in which the telescopes are all identical) can form pixellated transformation-optics (PTO) devices [16, 17].

A Gabor superlens comprises inhomogeneous GCLAs (in which the telescopes are not identical): the two lenslet arrays are confocal, but the separation between neighbouring lenslets is different in the two arrays. (Note that they are not rotated with respect to each other, which would result in a Moiré magnifier [18].)

It is important to understand that GCLAs (including Gabor superlenses) suffer from a number of imperfections in addition to those associated with pixel visibility and diffraction mentioned above. Perhaps the most noticeable of these imperfections results from light incident from certain directions entering through the first lens of one telescope and exiting through the second lens of another telescope. Such light forms additional images [19]. In principle, it can be absorbed with baffles separating the individual telescopes, but at the price of reduced transmission. The associated field-of-view limitation and reduction in transmission coefficient have been quantified for simple geometries [20]. Other imperfections are related to the limited imaging quality of the simple lenses used in designs that can be easily produced, each lens simply consisting of a (in the simplest case spherical) bump on the surface of a plastic sheet. These imperfections are currently the topic of ongoing optical-engineering efforts [21, 22].

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**Fig. 1.** (a) Rays passing through a generalised thin lens (glens). It can be seen that the focal points, $F^−$ and $F^+$, are located at different distances from the glens. The glens is indicated by a grey line with asymmetric triangles on both ends, a modification of the traditional symmetric triangles at the ends of a thin lens to reflect the asymmetry between the spaces on either side. The optical axis is shown as a dash-dotted line. It coincides with the axis of a Cartesian coordinate system, with corresponding coordinate $a$, such that the glens lies in the plane $a = 0$. We use the same coordinate system for the spaces on both sides of the glens. The triangles at the ends of the line indicating the glens are drawn on the side of the glens facing in the positive $a$ direction. Another way to indicate the sense of the optical axis is to draw next to the glens a ‘−’ on the side of negative space and a ‘+’ on the side of positive space. $F^−$ and $F^+$ are the focal points in the glens’s positive (positive) and negative (negative) space; $f^−$ and $f^+$ are the corresponding $a$ coordinates, the focal lengths; the focal planes are shown as dotted vertical lines. A number of rays (solid red lines) are shown, namely the principal rays through the positive and negative focal points, the ray through the glens centre, $P$, where the optical axis intersects the glens, and the undeviated ray between the object-image pair $Q^−$ and $Q^+$, which passes through the nodal point, $N$. (b) Glen Clova, one of the Scottish valleys also known as glens.
3. THE NODAL POINT

The behaviour of a thin lens is fully determined by the positions of the plane of the lens, its optical axis, and the focal points. As a thin lens does not offset light rays upon transmission, any light ray intersecting the thin lens enters the lens and leaves it from the same position in the lens plane. The optical axis, which is perpendicular to the lens plane, is the axis of cylindrical symmetry of the lens. Using imaging language — before transmission through the lens, light rays are in object space, afterwards they are in image space —, the focal points can be defined as those points where light rays that are parallel to the optical axis in object space intersect in image space (the image-sided focal point), and where light rays that are parallel to the optical axis in image space intersect in object space (the object-sided focal point). They are located on the optical axis on both sides of the lens, a distance \( f \) (the focal length) from the lens plane. The lens centre, where the optical axis intersects the lens plane, has the property that any light ray incident on it passes straight through it. This property is very useful, for example when drawing a principal-ray diagram to construct the position of the image of a given object position.

The behaviour of a glens is also fully determined by the positions of the glens plane, the optical axis, and the focal points [1]. It is similar to a lens in the following ways:

1. A glens also does not offset light rays upon transmission, and so any light ray intersecting a glens is incident on, and leaves from, the same position in the glens plane.

2. A glens also has an optical axis which is perpendicular to the glens plane and which is the axis of cylindrical symmetry.

3. A glens also has focal points where light rays that are parallel to the optical axis in object (image) space intersect in object (image) space, and these are located on the optical axis.

A glens also differs from a lens in a number of ways:

1. The focal points are, in general, located at different distances from the glens, and might even lie on the same side of the glens.

2. Light rays through the centre of a glens in general change direction. The ray diagram shown in Fig. 1(a) illustrates a few of these properties.

Fig. 1(a) also introduces a number of conventions for glenses. Firstly, we place on the optical axis one of the axes of a Cartesian coordinate system. We call the associated coordinate \( a \); the origin is chosen such that the glens lies in the \( a = 0 \) plane. This allows us to distinguish between light rays that approach the glens from negative space, i.e. from the direction of negative \( a \) values, from those that approach from positive space, i.e. from the direction of positive \( a \). This, in turn, allows us to distinguish between the focal lengths in negative space and in positive space, which is required here as these are, in general, different (unlike in the case of a lens). The focal lengths in negative and positive space, \( f^- \) and \( f^+ \), are defined as the \( a \) coordinates of the corresponding focal points, \( F^- \) and \( F^+ \). (Throughout this paper, we continue to label positions and distances that refer to negative space with a superscript ‘−’, those that refer to positive space with a superscript ‘+’.) The glens plane is indicated by a line with triangles at the ends. Unlike the triangles at the end of the lines indicating a thin-lens plane, the triangles indicating a glens are drawn asymmetrically, on the positive side of the glens, representing the broken symmetry between the spaces on either side. The negative and positive sides of a glens can also be indicated by placing a ‘−’ and a ‘+’ on the relevant sides of the glens. Note that either space can be object or image space; for example, if light is incident from negative space, then negative space is object space and positive space is image space.

We stress that our definition of the focal lengths is different from the standard definition; if the focal lengths according to the standard definition are \( f \) on the positive side and \( g \) on the negative side, then \( f^+ = f \) and \( f^- = -g \). The reason behind this redefinition is that it uses the same coordinate system on both sides of the glens, which is convenient for our purposes in that any distance which simultaneously refers to negative space and positive space can be defined unambiguously. We will see below that the nodal distance, the \( a \) coordinate of the nodal point, is such a distance.

All rays shown in Fig. 1(a) are involved in the imaging between two conjugate positions, \( Q^- \) and \( Q^+ \). As a glens images, by definition, each object-space point into a corresponding image-space position, any light ray that intersects the object position in object space intersects the image position in image space. To construct the image position, it is sufficient to construct the intersection of two such rays, and in Fig. 1(a) this was done with the ray that passes through the object-sided focal point and the ray that is incident parallel to the optical axis, and which passes through the image-sided focal point. The trajectories of additional rays involved in the imaging between \( Q^- \) and \( Q^+ \) can be constructed as two straight-line segments, one from \( Q^- \) to any point in the glens plane, the second from the same point in the glens plane to \( Q^+ \). A glance at the figure reveals that the ray through the glens centre, \( P \), does not pass through the glens undeviated, reducing the usefulness of \( P \) in ray diagrams. But Fig. 1(a) also shows that there exists another ray that passes through both \( Q^- \) and \( Q^+ \) and which passes through the glens plane undeviated, but this ray does not pass through the glens centre but instead intersects the optical axis at a different position, \( N \).

With the help of Fig. 2, we can demonstrate that any light ray incident on \( N \) — from any direction — passes through the glens undeviated. For reasons that will become clear, we call \( N \) the nodal point of the glens. (It is, of course, also possible to
demonstrate the existence of the nodal point from the mapping equations derived in Ref. [1]. Consider a light ray, marked 'Y', incident from negative space with an arbitrary direction d through the focal point F. After transmission through the glens, its direction is parallel to the optical axis. We call the point where it intersects the positive focal plane I. Now consider a second light ray, marked 'Z', incident with the same direction d and aiming at the point I. As the two rays are parallel in negative space, they intersect in a point in the positive-space focal plane. This point is I, and the second light ray is already intersecting it in negative space, i.e. before transmission through the glens. In order to intersect it also after transmission through the glens, its direction therefore cannot be altered by transmission through the glens.

It can easily be seen that the second, undeviated, ray always intersects the optical axis at the same position, N, which is, as the name choice suggests, the position of the nodal point. Because of the identity of the triangles F−SP and NIF+, N lies a distance \( f^+ \) from \( F^+ \), and a distance \( f^- \) from \( F^- \).

\[
n = f^+ + f^- \tag{1}
\]

from the glens centre, P. We call \( n \) the nodal distance of the glens. In the special case of a thin lens, for which \( f^- = -f^+ \), the nodal distance is \( n = 0 \) and so the nodal point, N, coincides with the lens centre. Eqn (1) shows that the position of N is independent of the direction d of the incident rays, which justifies our earlier assertion that any ray incident on N with any direction d passes through the glens undeviated.

Finally, we put our results in the context of the standard theory of ideal imaging systems, which allows the behaviour of an ideal imaging system to be determined by the locations of a number of cardinal points [23].

Firstly, above we reviewed the property of a glens that it does not offset light rays upon transmission, and so any light ray passing through the glens leaves it from the same point it is incident on. This is specifically the case for any light ray parallel to the optical axis, which means [23] that the glens plane is a principal plane. There are actually two principal planes, each corresponding to light rays incident from one side of the imaging system (here the glens), and it is immediately clear that both those principal planes coincide in the glens plane. The principal points, which are defined as the points the two principal planes intersect the optical axis, coincide with the glens centre, which is therefore the glens’s principal point, which we had presciently named P.

Secondly, the nodal points of an ideal imaging system are defined as the two points on the optical axis with the property that any light ray incident on one emerges from the optical system through the other, with the same light-ray direction. In the case of a glens, the two nodal points coincide in the point N, so it is indeed the glens’s nodal point.

4. THE MAPPING BETWEEN OBJECT AND IMAGE SPACE

We now use the properties of the nodal point to derive a simple equation for the mapping between object and image space. Fig. 3 shows the geometry. It shows two light rays involved in the imaging between a pair of conjugate points, Q (object space) and Q (image space). Using standard notation, we write quantities that refer to object space as unprimed, those that refer to image space as primed. Both light rays pass through Q and Q', addi-
is easily transformed into the form
\[ \frac{f}{a} + \frac{f'}{a'} = 1. \]  
(7)

In analogy to its lens equivalent, we call this equation the glens equation. Note that it has retained the form of the third equation in Eqns (10) in Ref. [1], despite the fact that we have redefined the sign of the \( a \) component of negative space positions (as in the term referring to negative space both numerator and denominator change sign).

3. When the change in the definition of the \( a \) components is taken into account, the \( b \) and \( c \) components of the equation are the same as the first two equations in Eqns (10) in Ref. [1].

4. The transverse (perpendicular to the axial direction) magnification of the mapping is
\[ M_T = \frac{\partial b'}{\partial b} = -\frac{f}{a - f}. \]  
(8)

(The same result is obtained by calculating \( \frac{\partial c'}{\partial c} \).) There is only one plane that is imaged with \( M_T = 1 \), namely the plane of the glens itself, the plane \( a = 0 \).

5. The longitudinal (in the axial direction) magnification of the mapping is
\[ M_L = \frac{\partial a'}{\partial a} = -\frac{f(n - f)}{(a - f)^2} = -\frac{ff'}{(a - f)^2} = \frac{f'}{f} M_T. \]  
(9)

In the lens case \( (f' = -f) \), \( M_L = M_T^2 \).

6. From the \( a \) component of Eqn (4) it is immediately clear that any object in the plane with \( a = n \) is imaged again into the same plane. We call this plane the nodal plane. The point where the nodal plane intersects the optical axis is, of course, the nodal point N. Substitution of \( a = n \) into Eqns (8) and (9) reveals that points in the nodal plane are imaged with the same transverse and longitudinal magnification, namely
\[ M_T = M_L = -\frac{f}{f'}. \]  
(10)

7. From the \( a \) component of Eqn (4) it can be seen that exchanging the order of two parallel glenses, one immediately in behind the other and sharing an optical axis, changes the mapping of the combination, unless both have the same nodal distance. If the object-sided focal lengths of the first and second glens are \( f_1 \) and \( f_2 \), and their nodal distances are \( n_1 \) and \( n_2 \), then the combination images an object position with longitudinal coordinate \( a \) to an image position with longitudinal coordinate
\[ a' = \frac{a(f_1 - n_1)(f_2 - n_2)}{f_1 f_2 - a(f_1 + f_2) + an_1}. \]  
(11)

Exchanging the order of the glenses leaves the numerator and the first two terms of the denominator unchanged, but the third term of the denominator changes from \( an_1 \) to \( an_2 \). If \( n_1 = n_2 \), this does not actually change the third term of the denominator. A calculation of the transverse coordinates of the image reveals that, provided \( n_1 = n_2 \), reversing the order of the glenses leaves all coordinates of the image unchanged.

It is also instructive to investigate the mapping in the limit when glenses become homogeneous, the imaging sheets studied in Ref. [24]. This can be visualised as picking a point on the surface of a glens, and then isotropically and infinitely magnifying the glens with the point on the surface as the centre of the magnification. The focal lengths then become infinite, and the nodal point N becomes infinitely distant. In this limit, \( f = a \to f \) and \( Q - N \to -N \), and so the mapping equation, Eqn (3), becomes
\[ Q' = Q - \frac{N}{f}. \]  
(12)

Comparison with Eqn (11) in Ref. [24] reveals that
\[ \frac{N}{f} = \left( \begin{array}{c} \delta_x \\ \delta_y \\ 1 - \eta \end{array} \right), \]  
(13)

where \( \delta_x, \delta_y, \) and \( \eta \) are parameters describing the imaging sheet.

Finally, we point out that Eqn (5) is in the form of a collineation (also known as a projective transform), the most general bijective (i.e. one-to-one and onto) transformation that takes lines to lines and planes to planes [25]. It is, of course, necessary that this transformation takes lines to lines as all object positions on any incident light ray must be imaged to corresponding image positions on the corresponding outgoing light ray light, and the trajectories of both the incident and outgoing rays are straight lines.

Translating two of the properties of glenses into the language of collineations allows us to identify the mapping more precisely. The first of these properties is that light rays exit a glens from the same positions where they entered it, but on the other side. This means that any point in the plane of the glens is imaged to itself, and therefore a fixed point of the collineation. In 2D, the straight line on which the fixed points lie would be called the axis of the collineation. The second of these properties is the existence of the nodal point, which lies on the straight line through any pair of object and image positions. In the language of collineations, the nodal point would be called the center of the collineation, and by virtue of possessing a centre the glens mapping is a central collineation [26]. The relationship between the axis and the centre of the collineation defines the subclasses of collineations [27] that describe the mapping due to thin lenses and glenses: In thin lenses, the centre lies on the axis of collineation, which makes the mapping an elation. In glenses, the centre does not in general lie on the axis of collineation, which makes the mapping a homology.

5. LIGHT-RAY-DIRECTION CHANGE UPON TRANSMISSION THROUGH A GLENS AND ITS IMPLEMENTATION IN DR TIM

We now derive equations describing the light-ray-direction change upon transmission through a glens. We write these equations in vector form to facilitate implementation in our ray-tracing software, DR TIM [28].

Fig. 4 sketches the trajectory of a light ray that is incident, with non-zero but otherwise arbitrary direction \( \mathbf{d} \), on an arbitrary point \( S \) on a glens. As before, we refer to the space from which the ray is incident as object space (unprimed), the other space as image space (primed).

The direction \( \mathbf{d}' \) of the light ray after transmission through the glens can be constructed using the fact that parallel incident
We have used this new capability to simulate the view through a glens. A light ray (solid red line) is incident from object space, with direction $\mathbf{d}$, at position $S$. If the light ray is incident from positive (negative) space, then object space is positive (negative) space and image space is negative (positive) space. The auxiliary light ray (dotted red line) that passes through the nodal point $N$ with the same direction intersects the image-sided focal plane at the same position, $I$, as the refracted light ray.

Light rays, after transmission through the glens, intersect in the image-sided focal plane. The direction of the redirected light ray is then

$$d' \propto I - S,$$  \hspace{1cm} (14)

where $I$ is the position vector corresponding to the point where the light ray intersects the image-sided focal plane, $I$, and $S$ is the position vector that corresponds to the position $S$. If $F'$ lies on the object-space side, then the vectors $I - S$ and $d'$ point in opposite directions. This happens when $f'$ and $d_a$, where

$$d_a = \mathbf{d} \cdot \hat{a}$$  \hspace{1cm} (15)

is the component of $\mathbf{d}$ in the direction of the optical axis, have opposite signs. Multiplying by the ratio of $d_a$ and $f'$ reverses the ray when this is the case, so we find

$$d' = \frac{d_a}{f'} (I - S).$$  \hspace{1cm} (16)

(Note that $f' \neq 0$ in all physically meaningful cases.)

To calculate $I$, we consider a principal ray that intersects $N$ (position vector $N$) and that is parallel to the incident ray. We can then write $I$ in the form

$$I = N + f \frac{d}{d_a}. $$  \hspace{1cm} (17)

Substitution into Eqn (16) gives

$$d' = -\frac{f'}{f} d + \frac{d_a}{f'} (N - S). $$  \hspace{1cm} (18)

6. RAYTRACING SIMULATIONS

We have used the equation describing the change of light-ray direction upon transmission through a glens, Eqn (18), to add the capability to simulate transmission through a glens to our custom raytracer Dr TIM [28]. (More precisely, we have used Eqns (21) and (22) (see A), which are versions of Eqn (18) that also work in the case of infinite focal lengths and nodal distance.) We have used this new capability to simulate the view through glenses, concentrating on phenomena associated with pseudo-sopic imaging and the location of the nodal point outside of the plane of the glens, which do not occur for lenses.

Fig. 5 simulates the view through a glens of objects in the glens’s nodal plane. Frames (b) and (d) in Fig. 5 show a collection of objects located close to the nodal plane of a glens, seen through that glens from different directions. The objects touch the nodal plane from behind; the vertex of the cone is located at the nodal point, $N$. The glens present in (b) and (d) is located a distance 2 (in units of floor-tile widths) in front of the nodal plane, with its $a$ axis pointing away from the camera in (b). Its focal lengths are $f^- = -2$ and $f^+ = 4$. In all frames, the virtual camera is focussed on the nodal plane.

Fig. 6 demonstrates what happens when the camera is located at the nodal point. If the camera is a pinhole camera, every light ray recorded by the camera is a light ray through the glens’s nodal point, and such light rays do not change direction when passing through the glens. This is why the view in frame (a), which shows a scene with the glens absent, looks the same as that in frame (b), in which the glens is present. Each image produced by the glens is therefore seen in the same direction as the corresponding object. Frames (c) and (d) illustrate that these images are formed at distances different from those of the objects. In frame (c), in which the virtual camera is focussed on the (front of the) objects in the scene, the objects are therefore blurred, but in frame (d), in which the camera is focussed on the

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**Fig. 4.** Light-ray-direction change upon transmission through a glens. A light ray (solid red line) is incident from object space, with direction $\mathbf{d}$, at position $S$. The auxiliary light ray (dotted red line) that passes through the nodal point $N$ with the same direction intersects the image-sided focal plane at the same position, $I$, as the refracted light ray.

**Fig. 5.** Ray-tracing simulations of objects located near the nodal plane of a glens, seen through that glens (b,d), and for comparison without the glens (a,c). (a) and (b) have been calculated for one position of the virtual camera, (c) and (d) for another. The sphere, the cone, and the cylinder touch the nodal plane from behind; the vertex of the cone is located at the nodal point, $N$. The glens present in (b) and (d) is located a distance 2 (in units of floor-tile widths) in front of the nodal plane, with its $a$ axis pointing away from the camera in (b). Its focal lengths are $f^- = -2$ and $f^+ = 4$. In all frames, the virtual camera is focussed on the nodal plane.

**Fig. 6.** Demonstrates what happens when the camera is located at the nodal point. If the camera is a pinhole camera, every light ray recorded by the camera is a light ray through the glens’s nodal point, and such light rays do not change direction when passing through the glens. This is why the view in frame (a), which shows a scene with the glens absent, looks the same as that in frame (b), in which the glens is present. Each image produced by the glens is therefore seen in the same direction as the corresponding object. Frames (c) and (d) illustrate that these images are formed at distances different from those of the objects. In frame (c), in which the virtual camera is focussed on the (front of the) objects in the scene, the objects are therefore blurred, but in frame (d), in which the camera is focussed on the
Finally, we simulate the view through a telescope formed by glenses that share a common focal point (the positive focal point of the first glens coincides with the negative focal point of the second) and a common nodal point (Fig. 7). Such a telescope changes the beam size without changing its propagation direction. This is forbidden by Liouville’s theorem as phase-space volume is not preserved, reflecting the fact that glenses are wave-optically forbidden. (Note that, conversely, a single homogeneous glens with \( f^- \rightarrow \infty \) such that \( f^- / f^+ = \text{const.} \) changes the beam angle without changing its size.) The mapping between object and image space created by such a telescope is an isotropic stretching by a factor \( M \), centred on N. When seen from N, any object is seen in the “true” direction, but a factor \( M \) further away. This is demonstrated in Fig. 8.

Fig. 7. Glens telescope comprising a pair of glenses that share a common focal point, namely the first glens’ positive-space focal point, \( F_1^+ \), and the second glens’ negative-space focal point, \( F_2^- \), and a common nodal point, N. The trajectories of a bundle of initially parallel rays (red lines) is shown. As it passes through the telescope from left to right, the ray bundle gets expanded without reducing its angle with the optical axis.

Fig. 8. Simulated view through the glens telescope shown in Fig. 7, which multiplies the distance of any light ray from N by a factor \( M = 3 \) without altering the ray’s direction. The camera is positioned at the glenses’ common nodal point, N. In (a), the camera is focussed on the plane where the objects actually are (here, \( z = 10 \)). In (b), it is focussed on a plane \( M \) times further away (\( z = 30 \)). In both frames, the central, circular, part of the view is seen through both glenses, which provide a sharp image in (b).

7. CONCLUSIONS

In this paper, we have defined glenses as idealised thin lenses, generalised such that the focal lengths on both sides can be different. Glenses can be approximately realised in the form of Gabor superlenses. We have shown that a glens possesses a nodal point that can be located anywhere on the optical axis. When formulated in terms of this nodal point, the equations describing the mapping between object and image space and the light-ray-direction change on transmission through a glens become particularly simple.

We intend to apply our findings to the construction of novel optical instruments, starting with a generalisation of the pixelated transformation-optics devices described in Ref. [17].

8. ACKNOWLEDGEMENTS

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A. REPRESENTATION OF GLENS PARAMETERS IN DR TIM

It can happen that both focal lengths become infinite, namely in the limit of homogeneous glenses \[24\], in which case the equations for the mapping, Eqn (3), and for the light-ray-direction change, Eqn (18), do not work. It would be possible to switch to alternative equations, but we took a different approach when programming glenses into Dr TIM, namely to formulate the glenses in terms of parameters that work irrespective of whether or not the focal lengths are finite or infinite.

The solution is to divide the quantities that go to \( \infty \) by a length, \( g \), such that they become finite, and describe the glens in terms of those dimensionless parameters, namely \( (f/g), (f'/g), (N/g) \). The length \( g \) itself can become infinite. In the case of glenses with finite focal lengths, we simply choose \( g = 1 \) (Dr TIM uses dimensionless units). In the case of glenses with infinite focal lengths, we choose \( g = n \), the nodal distance.

Formulated in terms of \( g \), the mapping equation, Eqn (3), becomes

\[
Q' = Q + \frac{a}{f/g - a/g} \left( \frac{Q}{g} - \frac{N}{g} \right),
\]

(19)
In the limit \( g \to \infty, a/g \to 0 \) for any finite value of \( a \) and \( Q/g \to 0 \) for any finite object position \( Q \), and so

\[
Q' = Q - \frac{aN}{f/g}.
\]  

The equation describing the light-ray-direction change, Eqn (18), can also be formulated in terms of \( g \). It becomes

\[
d' = -\frac{f/g - d}{f/g} + d_a \left( \frac{N}{g} - \frac{S}{g} \right).
\]

For a finite intersection point \( S \), this simplifies to

\[
d' = -\frac{f/g - d}{f/g} + d_a \frac{N}{g}.
\]

REFERENCES