Moderní experimentální metody Rentgenová a elektronová spektroskopie I

Spektroskopie absorpční hrany rtg záření

- Principy, experimentální realizace
- Metody: XANES, EXAFS
- Postupy vyhodnocení dat, příklady
- Magnetismus XMCD
- RIXS

Moderní experimentální metody Rentgenová a elektronová spektroskopie II

Anomální rtg difrakce

- Kramersovy-Kronigovy relace
- Anomální difrakce principy, použití
- DAFS

Flurescenční spektroskopie

- Spektroskopie charakteristického záření (XRF, EDS, WDS)
- Kvalitativní a kvantitativní analýza
- TRXRF
- Fluorescence ve stojaté vlně (SW-XRF) rtg reflexe, difrakce

Moderní experimentální metody Rentgenová a elektronová spektroskopie III

Fotoelektronová spektroskopie

- Fotoelektronová spektroskopie (XPS) a spektroskopie Augerových elektronů (AES)
- Úhlově rozlišená fotoelektronová spektroskopie (ARPES)
- Experimentální aspekty
- Zdroje: ARPES, ARUPS
- Detektory
- Příprava vzorků

Absorpce rtg záření

Ionizace hluboké elektronové hladiny – absorpční hrana.



Závislost absorpce na vlnové délce pro olovo (Z=82).

Empirická závislost absorpce na energii mimo absorpční hranu:

$$\alpha(\epsilon) = \frac{a}{\epsilon^3} + \frac{b}{\epsilon^4} \;,$$

Rentgenová absorpční spektroskopie jemné struktury X-ray absorption fine structure – XAFS Extended x-ray absorption fine structure – EXAFS Near-edge x-ray absorption fine structure – NEXAFS X-ray absorption near-edge structure – XANES



Incident Energy (eV)

Rentgenová absorpční spektroskopie jemné struktury Uspořádání experimentu ESRF BM26A



Rentgenová absorpční spektroskopie jemné struktury Uspořádání experimentu



Rentgenová absorpční spektroskopie jemné struktury



$$\mu_{\rm tot}(E) = \mu_{\chi}(E) + \mu_{\rm other}(E)$$

Rentgenová absorpční spektroskopie jemné struktury

Eq. 4.2 has several interesting limits that are common for real XAFS measurements. First, there is the *thin sample limit*, for which $\mu t \ll 1$. The $1 - e^{-\mu t}$ term then becomes $\approx [\mu_{tot}(E)/\sin\theta + \mu_{tot}(E_f)/\sin\phi] t$ (by a Taylor series expansion), which cancels the denominator, so that

$$I_f \approx I_0 \frac{\epsilon \Delta \Omega}{4\pi} \mu_{\chi}(E) t \tag{4.4}$$

Alternatively, there is the *thick*, *dilute sample limit*, for which $\mu t \gg 1$ and $\mu_{\chi} \ll \mu_{\text{other}}$. Now the exponential term goes to 0, so that

$$I_f = I_0 \frac{\epsilon \Delta \Omega}{4\pi} \frac{\mu_{\chi}(E)}{\mu_{\text{tot}}(E) / \sin \theta + \mu_{\text{tot}}(E_f) / \sin \phi}.$$
(4.5)

Finally, One way to reduce these self-absorption effects for thick, concentrated samples is to rotate the sample so that the sample is normal to the incident beam. With $\phi \to 0$ or the grazing exit limit, $\mu_{tot}(E_f)/\sin\phi \gg \mu_{tot}(E)/\sin\theta$, giving

$$I_f \approx I_0 \frac{\epsilon \Delta \Omega}{4\pi} \frac{\mu_{\chi}(E)}{\mu_{\text{tot}}(E_f) / \sin \phi}$$
(4.7)

XANES

Profil absorpční hrany závisí na chemickém stavu atomu. Nutný kvantově-mechanický výpočet.

Přechod z hluboké hladiny do volných stavů nad Fermiho mezí.

Příklad pro železo v různých chemických stavech.



Jemná struktura okolí absorpční hrany – určení poloh nejbližších atomů.



Vlnový vektor excitovaného elektronu

$$k_{\rm e}^2 = \frac{2m_{\rm e}}{\hbar^2} (\epsilon - \epsilon_0)$$

Příspěvek jednoho atomu ve vzdálenosti R

$$\chi(k) = \frac{f(k)}{kR^2} \sin[2kR + \delta(k)]$$

Celková amplitude středovaná přes polohy a typ sousedů

Závislost úměrná R⁻² – citlivé maximálně na třetí nejbližší sousedy.

Příspěvek jednoho atomu ve vzdálenosti R

$$\chi(k) = \sum_{j} \frac{N_{j} e^{-2k^{2} \sigma_{j}^{2}} e^{-2R_{j}/\lambda(k)} f_{j}(k)}{kR_{j}^{2}} \sin[2kR_{j} + \delta_{j}(k)]$$

Celková amplitude středovaná přes polohy a typ sousedů

Závislost úměrná R⁻² – citlivé maximálně na třetí nejbližší sousedy.







Normovaný koeficient absorpce pro měď. Měření čárkovaně, simulace plnou čarou.



Fe-O data



EXAFS + XANES analytický software

- FEFF
- FDMNES
- Demeter (Bruce Ravel):
 - Athena
 - Artemis
 - Hephaestus



EXAFS

Mn in BiTe





Mn in BiTe





Absorpce na hladinách s orbitálním momentem

Povrchově citlivé, Prvkově citlivé



Asymmetry $A=(I_+-I_-)/(I_++I_-)$

Absorpce na hladinách s nenulovým orbitálním momentem (vyšší než K)

Spin and Orbital Moments: X-Ray Magnetic Circular Dichroism





X-rays absorption and X-ray magnetic circular dichroism (XMCD)



Two step model:

- 1) spin-polarized photoelectrons are created by using circularly polarized x-rays
- 2) these polarized photoelectrons are used to analyze the spin-split valence density of states, thus the valence band acts as a spin-sensitive detector.







 $\underline{\varepsilon}_{\alpha}$ is the polarization

dependent electric vector

Photon-electron interaction: dipole approximation H_{int}

$$H_{int}(0,t) \approx \underline{r} \cdot \underline{\varepsilon}_{q} \left[a_{k} \exp(-i\omega_{k}t) + c.c. \right]$$

The dipole operator $P_q^{1} = \underline{r} \cdot \underline{\varepsilon}_q$ can be written in terms of Racah's tensor operators (where $Y_{1,m}$ are the spherical harmonics)

$$P_{\pm 1}^{(1)} = \mp \frac{1}{\sqrt{2}} (x \pm iy) = rC_{\pm 1}^{(1)} = r\sqrt{\frac{4\pi}{3}}Y_{1,\pm 1}$$

$$P_{\pm 1}^{(1)} = \mp \frac{1}{\sqrt{2}} (x \pm iy) = rC_{\pm 1}^{(1)} = r\sqrt{\frac{4\pi}{3}}Y_{1,\pm 1}$$

$$P_{0}^{(1)} = z = rC_{0}^{(1)} = r\sqrt{\frac{4\pi}{3}}Y_{1,0}$$

The photon absorption generates a transition from an initial core level to a final level close to the Fermi level (note that the spin is not affected)

$$|i\rangle = R_i(r) |c, m_c; s = \frac{1}{2}, m_s\rangle \qquad \longrightarrow \qquad |f\rangle = R_f(r) |l, m_l; s = \frac{1}{2}, m_s\rangle$$

Table 1.3: Electric dipole matrix elements $\langle l, m_l | C_q^{(1)} | c, m_c \rangle$ in the one electron model. The matrix elements are non-vanishing when c = l - 1 (left column) or c = l + 1 (right column), and when $m_c + q = m_l$. q denotes the state of polarization of the photons which mix the states $|l, m_l\rangle$ and $|c, m_c\rangle$.

Remember: things become easily complicate. For example, if spin-orbit interaction is not negligible the (L, m_L , S, m_S) is not the good basis and you have to use the (L, S, J, m_J) basis



X-rays absorption and X-ray magnetic circular dichroism (XMCD)







X-rays absorption and X-ray magnetic circular dichroism (XMCD)





Fig. 9.14. Polarization dependent transition intensities in a one-electron model from spin-orbit and exchange split p core states $|j, m_j\rangle$ to spin-up $(m_s = +1/2) d$ valence orbitals (Table A.2), assumed to be split by the exchange interaction. The listed intensities each need to be divided by 90 to get the proper absolute values in units of \mathcal{AR}^2 . We have chosen the z-axis as the spin quantization axis and the transition intensities are for circular polarization with $k \parallel z$ and angular momenta q = +1(labeled +) and q = -1 (labeled -) and for linear polarization with $E \parallel z$ (labeled q = 0). We have assumed a splitting of the p states by the exchange interaction, lifting the degeneracy in m_j . Note that this causes an opposite order of m_j states for $p_{3/2}$, l + s and $p_{1/2}$, l - s because of the opposite sign of s



Generace polarizovaného rtg záření



Generace polarizovaného rtg záření



Lineární polarizátor pro rtg záření





 $P = \cos^2 2\theta$

P = 1

Case (1): P-Polarisation in plane of scattering, Case (2): S-Polarisation perpendicular to plane of scattering,

"Brewsterův úhel" = 45 stupňů



Generace polarizovaného rtg záření

Difrakční křivka Si 004, Cu K alfa, 2theta=69.5 deg



Generace polarizovaného rtg záření

"čtvrtvlnná destička"



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Orbital and spin sum rules in x-ray magnetic circular dichroism

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Electron yield $Y_i(\hbar\omega) \propto \hbar\omega\sigma_i(\hbar\omega)$,

$$\frac{\Delta A_{L_3} + \Delta A_{L_2}}{A_t} = \frac{\int_{L_3} \sigma_M(\hbar\omega) d\omega + \int_{L_2} \sigma_M(\hbar\omega) d\omega}{\int_{L_{2,3}} [\sigma_0(\hbar\omega) + \sigma_+(\hbar\omega) + \sigma_-(\hbar\omega)] d\omega}$$
$$= \frac{\langle L_z \rangle}{2n_b}$$
(8)

$$\frac{\Delta A_{L_3} - 2\Delta A_{L_2}}{A_t} = \frac{\int_{L_3} \sigma_M(\hbar\omega) d\omega - 2\int_{L_2} \sigma_M(\hbar\omega) d\omega}{\int_{L_{2,3}} [\sigma_0(\hbar\omega) + \sigma_+(\hbar\omega) + \sigma_-(\hbar\omega)] d\omega}$$
$$= \frac{2}{3n_h} \langle S_Z \rangle + \frac{7}{3n_h} \langle T_Z \rangle , \qquad (9)$$



FIG. 6. Fe, Co, and Ni XMCD spectra shown normalized to constant L_3 peak height. The features labeled B are due to multiple initial-state configurations. The constant negative intensity between B and the L_2 peak in the Ni and Co spectra is due to diffuse magnetism.

$$\langle L_z \rangle = -2n_{\rm h} \cdot \frac{\int_{\rm L_3 + L_2} [I^{\dagger\dagger} - I^{\dagger\downarrow}] d\omega}{\int_{\rm L_3 + L_2} [I^{\dagger\dagger} + I^0 + I^{\dagger\downarrow}] d\omega}$$

$$\langle S_{z} \rangle = -\frac{3}{2} n_{\mathrm{h}} \cdot \frac{\int_{\mathrm{L}_{3}} [I^{\dagger\dagger} - I^{\dagger\downarrow}] \mathrm{d}\omega - 2 \int_{\mathrm{L}_{2}} [I^{\dagger\dagger} - I^{\dagger\downarrow}] \mathrm{d}\omega}{\int_{\mathrm{L}_{3} + \mathrm{L}_{2}} [I^{\dagger\dagger} + I^{0} + I^{\dagger\downarrow}] \mathrm{d}\omega}$$

The final spectrum is the sum of 20 individual (F/I_0) sets for each side, with the X-ray photon angular momentum either parallel $(I^{\uparrow\uparrow})$ or antiparallel, $(I^{\uparrow\downarrow})$ to that of the C(II) majority 3d spin. Each of the averaged spectra were first normalized to

Here $n_{\rm h}$ represents the number of 3d vacancies in the metal ion, ω is the X-ray frequency, and I^0 refers to the absorption for X-rays linearly polarized along the magnetization direction; as is customary, we approximate this by $(I^{\uparrow\uparrow} + I^{\downarrow\downarrow})/2$. It is common in the XMCD literature^{1,7,15} to refer to the first integral in the numerator of eq 2 as "A", the second integral as "B", and the denominator as "3C". The necessary integrals are illustrated in Figure 1. One can thus re-express the sum rules in the following manner, in units of \hbar/a tom:

$$\langle L_z \rangle = -n_{\rm h} \frac{2(A+B)}{3C} \qquad \langle S_z \rangle = -n_{\rm h} \frac{A-2B}{2C} \qquad (3)$$



XMCD Mn in Bi2Se3



FIG. 4. (Color online) XMCD in total electron yield at the Mn $L_{2,3}$ -edges. (a) X-ray aborption spectra measured for opposite helicities of the circularly polarized incident light at 5 K with a out-of-plane applied magnetic field of 3 T. A sketch of the experimental geometry is also shown. (b) Corresponding XMCD difference spectrum. The inset shows XMCD measurements for different applied magnetic fields. (c) Temperature dependence of the remanent XMCD for a Mn concentration of x = 0.04. Inset: Detailed temperature dependence for various Mn concentrations.



FIG. 3. (Color online) (a) Hysteresis loops measured at a temperature of 4.2 K (squares) and 7 K (triangles), respectively. The applied magnetic field is perpendicular to the (111) sample surface (see sketch). Inset: A zoom-in around zero magnetic field showing hysteresis at 4.2 K and a paramagnetic state at 7 K. (b) Modified normalized Arrott plots at various temperatures from which a Curie temperature of 5.5 K is deduced. Between 4.25 and 6.75 K, data are shown for increments of $\delta T = 0.25$ K.

Rezonanční neelastický rtg rozptyl – rezonanční "Ramanův" rozptyl Výhoda Rtg proti VIS – směrové rozlišení, možnost pokrytí celé Brillouinovy zóny Nevýhoda – slabší jevy, experimentální obtížnost energiového rozlišení





Elementary excitations in condensed matter systems that can be measured by RIXS. The indicated energy scales are the ones relevant for transition metal oxides.





Cu L3 RIXS results of La2CuO4. The peak dispersion curve was extracted from the raw data (left panel) corresponding to the given dots in the Brillouin zone.

L. Braicovich (a), L.J.P. Ament (b), V. Bisogni (c), F. Forte (b,d), C. Aruta (e), G. Balestrino (f), N.B. Brookes (c), G.M. De Luca (e), P.G. Medaglia (f), F. Miletto Granozio (e), M. Radovic (e), M. Salluzzo (e), J. van den Brink (b,g), and G. Ghiringhelli (a), arXiv:0807:1140v1, (2008).



$$F(\Omega,\omega) = \sum_{f} \left| \sum_{n} \frac{\langle f | T_2 | n \rangle \langle n | T_1 | g \rangle}{E_g - E_n + \Omega - i \frac{\Gamma_{INT}}{2}} \right|^2 * \frac{\frac{\Gamma_{FIN}}{2\pi}}{(E_g - E_f + \Omega - \omega)^2 + \frac{\Gamma_{FIN}^2}{4}}$$

ESRF

A high resolution soft X-ray RIXS spectrometer reaching a combined resolving power E/ΔE better than 30,000. The spectrometer features an 11m long scattering arm capable of rotating over 100° without breaking vacuum, and a full in-vacuum 4-circle sample goniometer. The instrument was designed and developed in collaboration with Giacomo Ghiringhelli & Lucio Braicovich from Politecnico di Milano.



ESRF



Bodové rtg detektory

- Ionizační
- Geiger-Müller vysoké napětí, Townsendova lavina
- Proporcionální nižší napětí, rozlišení > 20 % (> 10³ eV)



Bodové rtg detektory

 Scintilační – rozlišení 5 až 10 % (10³ eV) Nal(Tl) – od 6keV účinnost téměř 100% Emise 410 nm – 3eV



Polovodičové – rozlišení až 2 % (10¹ - 10² eV)
Si, Ge



Polovodičové rtg detektory



Lineární rtg detektory

- Plynové ionizační, malý dynamický rozsah (50 kcps),
- Polovodičové velký dynamický rozsah





Plošné rtg detektory

- Plynové ionizační
- CCD rozlišení cca (50µm)²
- Film rozlišení cca (1µm)², nelineární, velmi pomalé
- Luminiscenční ("Image plate") rozlišení cca (10μm)², velké plochy rychlost 1 snímek cca 3 až 10 minut.
- Scintilační rozlišení až cca (3µm)², malá kvantová výtěžnost (tenký scintilátor)
- Polovodičové



Rtg detektory

Table 4-2. Properties of common x-ray detectors; ΔE is measured as FWHM.

Detector	Energy range (keV)	∆ <i>E/E</i> at 5.9 keV (%)	Dead time/event (µs)	Maximum count rate (s ⁻¹)
Gas ionization (current mode)	0.2–50	n/a	n/a	10 ^{11a}
Gas proportional	0.2-50	15	0.2	106
Multiwire and microstrip proportional	3–50	20	0.2	10 ⁶ /mm ²
Scintillation [NaI(Tl)]	3-10,000	40	0.25	2×10^{6}
Energy-resolving semiconductor	1-10,000	3	0.5–30	2×10^5
Surface-barrier (current mode)	0.1-20	n/a	n/a	108
Avalanche photodiode	0.1–50	20	0.001	108
CCD	0.1 - 70	n/a	n/a	n/a
Superconducting	0.1–4	< 0.5	100	5×10^3
Image plate	4-80	n/a	n/a	n/a

^a Maximum count rate is limited by space-charge effects to around 10¹¹ photons/s per cm³.